

Precision Observables $M_W, (g - 2)_\mu$

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main refs: [Heinemeyer, Hollik, DS, Weber, Weiglein, '06] (M_W)
[DS '06 (soon...)] (review on $(g - 2)_\mu$ and SUSY)

Outline

1 Motivation

2 M_W

3 $(g - 2)_\mu$

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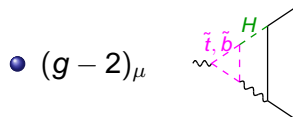
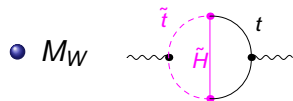
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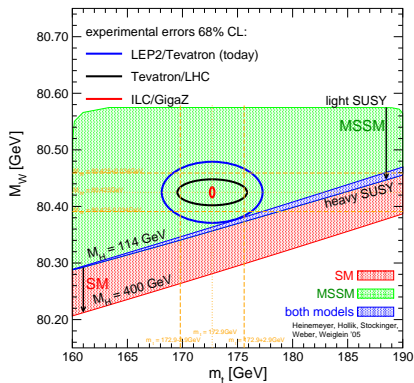
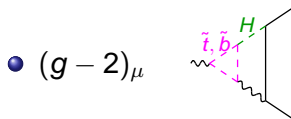
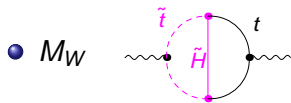
Precision Observables

Quantum effects from SUSY particles can have an influence on **precision observables**



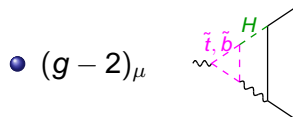
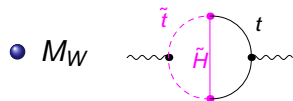
Precision Observables

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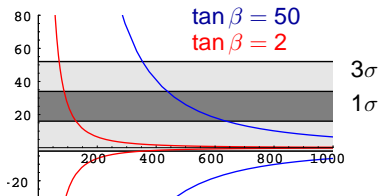
Precision Observables

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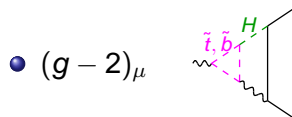
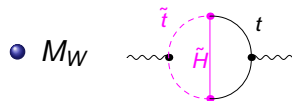
$$a_\mu(\text{exp} - \text{SM}) = 25(9) \times 10^{-10}$$

$$a_\mu(\text{SUSY})$$



Precision Observables

Quantum effects from SUSY particles can have an influence on **precision observables**



MSSM tends to agree better with experimental data than SM

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Muon-decay, M_W , and $\Delta\rho$

Muon lifetime ($\leftrightarrow G_\mu$) related to M_W :

$$M_W^2 \left(1 - \frac{M_W^2}{M_Z^2} \right) = \frac{\pi\alpha}{\sqrt{2}G_\mu} (1 + \Delta r)$$

\Rightarrow Prediction for M_W in terms of M_Z , α , G_μ , $\Delta r(m_t, M_H, m_{\tilde{t}}, \dots)$

Loop calculation:

$$\Delta r = -\frac{c_W^2}{s_W^2} \Delta\rho + \text{rest}$$

$\Delta\rho$: breaking of isospin invariance $\rightarrow t - b$ -mass splitting ...

Muon-decay: Status

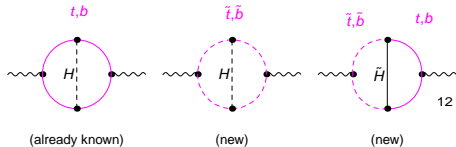
SM: $\Delta\rho$: 3-loop, 4-loop
rest: complete 2-loop

MSSM: $\Delta\rho$: 2-loop [Haestier,Heinemeyer,DS,Weiglein'05]
rest: 1-loop complete [Heinemeyer,Hollik,DS,Weber,Weiglein'06]

↓
new results

$\Delta\rho$ at 2-loop order

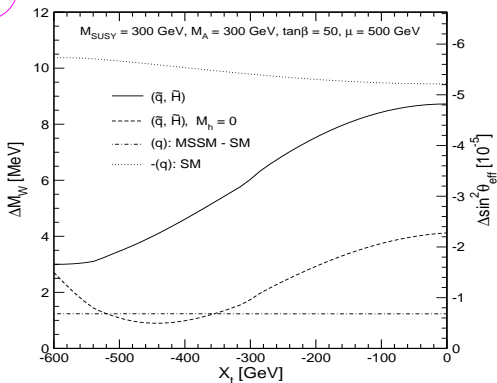
New result on Yukawa-enhanced corrections of $\mathcal{O}(\alpha_{t,b}^2)$



[Haestier, Heinemeyer, DS, Weiglein'05]

induced shift in M_W as
function of stop mixing: up to
 $\Delta M_W \approx 8$ MeV,

SUSY loops can be as large
as SM quark loops



New result for M_W in the MSSM

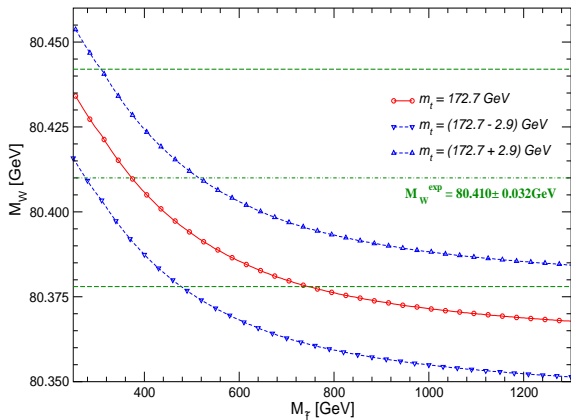
[Heinemeyer,Hollik,DS,Weber,Weiglein '06]

- Complete 1-loop calculation of Δr with complex parameters
- Incorporation of 2-loop calculation of $\Delta\rho$
[Haestier,Heinemeyer,DS,Weiglein'05]
- Incorporation of higher-order SM contributions

$$\Delta r^{\text{MSSM}} = \Delta r^{\text{SM}} + \Delta r_{1\text{-loop}}^{\text{MSSM-SM}} - \frac{C_W^2}{S_W^2} \Delta\rho_{2\text{-loop}}^{\text{MSSM-SM}}$$

Best available prediction of M_W in the MSSM \rightarrow implemented in computer program

Compare MSSM prediction for M_W with experimental result



[Heinemeyer, Hollik, DS, Weber, Weiglein '06]

$$\tan \beta = 10$$

$$A_{t,b} = 2M_{\tilde{f}}$$

$$\mu = 300 \text{ GeV}$$

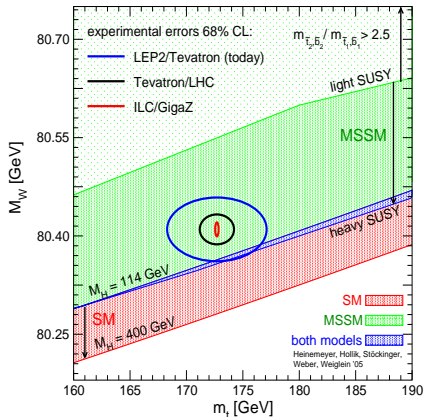
$$m_{\tilde{g}} = 300 \text{ GeV}$$

$$M_A = 300 \text{ GeV}$$

$$M_2 = 300 \text{ GeV}$$

Preference for light SUSY scale $M_{\tilde{f}} = M_{\tilde{t}, \tilde{b}}$

Compare MSSM prediction for M_W with experimental result



[Heinemeyer, Hollik, DS, Weber, Weiglein '06]

Scan of all SUSY parameters

Preference for MSSM over SM

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$(g - 2)_\mu$: Status

- BNL experiment: finalized, very stable development
- SM-theory: several errors fixed, $e^+ e^-$ data preferred

$$a_\mu(\text{exp} - \text{SM}) = (23.9 \pm 9.9) \times 10^{-10}$$

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$$a_\mu(\text{exp} - \text{SM}) = (23.9 \pm 9.9) \times 10^{-10}$$

Historical comparison: 1978, CERN experiment:

$$a_\mu(\text{CERN} - \text{SM}, \text{without had}) = (720 \pm 85) \times 10^{-10}$$

$$a_\mu(\text{SM}, \text{had}) = (667 \pm 81) \times 10^{-10}$$

Established had contributions, agreement with SM

$(g - 2)_\mu$: Status

- BNL experiment: finalized, very stable development
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$$a_\mu(\text{exp} - \text{SM}) = (23.9 \pm 9.9) \times 10^{-10}$$

Now: compare with weak contributions only:

$$a_\mu(\text{BNL} - \text{SM}, \text{without weak}) = (39.3 \pm 9.9) \times 10^{-10}$$

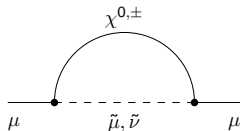
$$a_\mu(\text{SM}, \text{weak}) = (15.4 \pm 2.2) \times 10^{-10}$$

existence of weak contributions, but disagreement with SM! \rightarrow SUSY?

Status of SUSY prediction

1-Loop

$$\propto \tan \beta$$



[Fayet '80],...

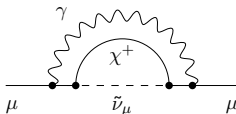
[Kosower et al '83],[Yuan et al '84],...

[Lopez et al '94],[Moroi '96]

complete

2-Loop logs

$$\propto \log \frac{M_{\text{SUSY}}}{m_\mu}$$

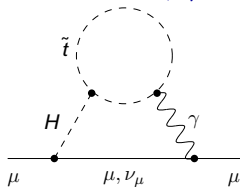


[Degrassi, Giudice '98]

complete

2-Loop non-log

$$\propto \tan \beta \mu m_t$$



[Chen, Geng'01][Arhib, Baek '02]

[Heinemeyer, DS, Weiglein '03]

[Heinemeyer, DS, Weiglein '04]

half

(no internal $\tilde{\mu}, \tilde{\nu}$)

SUSY prediction

Implementation of 1-Loop and leading 2-Loop straightforward:

1-Loop:
$$a_\mu^{\chi^0} = \frac{m_\mu}{16\pi^2} \sum_{i,m} \left\{ -\frac{m_\mu}{12m_{\tilde{\mu}m}^2} (|n_{im}^L|^2 + |n_{im}^R|^2) F_1^N(x_{im}) + \frac{m_{\chi_i^0}}{3m_{\tilde{\mu}m}^2} \text{Re}[n_{im}^L n_{im}^R] F_2^N(x_{im}) \right\},$$

$$a_\mu^{\chi^\pm} = \frac{m_\mu}{16\pi^2} \sum_k \left\{ \frac{m_\mu}{12m_{\tilde{\nu}\mu}^2} (|c_k^L|^2 + |c_k^R|^2) F_1^C(x_k) + \frac{2m_{\chi_k^\pm}}{3m_{\tilde{\nu}\mu}^2} \text{Re}[c_k^L c_k^R] F_2^C(x_k) \right\},$$

$$n_{im}^L = \frac{1}{\sqrt{2}} (g_1 N_{i1} + g_2 N_{i2}) U_{m1}^{\tilde{\mu}*} - y_\mu N_{i3} U_{m2}^{\tilde{\mu}*},$$

$$n_{im}^R = \sqrt{2} g_1 N_{i1} U_{m2}^{\tilde{\mu}} + y_\mu N_{i3} U_{m1}^{\tilde{\mu}},$$

$$c_k^L = -g_2 V_{k1},$$

$$c_k^R = y_\mu U_{k2}. F_1^N(x) = \frac{2}{(1-x)^2} [1 - 6x + 3x^2 + 2x^3 - 6x^2 \log x],$$

$$F_2^N(x) = \frac{3}{(1-x)^3} [1 - x^2 + 2x \log x],$$

$$F_1^C(x) = \frac{2}{(1-x)^2} [2 + 3x - 6x^2 + x^3 + 6x \log x],$$

$$F_2^C(x) = \frac{3}{(1-x)^3} [-3 + 4x - x^2 - 2 \log x],$$

SUSY prediction

Implementation of 1-Loop and leading 2-Loop straightforward:

2-Loop:
$$a_\mu^{\text{logs}} = -\frac{4\alpha}{\pi} \log \frac{M_{\text{SUSY}}}{m_\mu} a_\mu^{\text{1-Loop}}$$

$$a_\mu^{(\chi\gamma H)} = \frac{\alpha^2 m_\mu^2}{8\pi^2 M_W^2 S_W^2} \sum_{k=1,2} \left[\text{Re}[\lambda_\mu^{A^0} \lambda_{\chi_k^+}^{A^0}] f_{\text{PS}}(m_{\chi_k^+}^2/M_{A^0}^2) + \sum_{S=H^0, H^0} \text{Re}[\lambda_\mu^S \lambda_{\chi_k^+}^S] f_S(m_{\chi_k^+}^2/M_S^2) \right],$$

$$a_\mu^{(\bar{l}\gamma H)} = \frac{\alpha^2 m_\mu^2}{8\pi^2 M_W^2 S_W^2} \sum_{\bar{l}=\bar{t}, \bar{b}, \bar{\tau}} \sum_{i=1,2} \left[\sum_{S=H^0, H^0} (N_C Q^2)_{\bar{l}} \text{Re}[\lambda_\mu^S \lambda_{\bar{l}}^S] f_{\bar{l}}(m_{\bar{l}}^2/M_S^2) \right].$$

$$\lambda_\mu^{\{H^0, H^0, A^0\}} = \left\{ -\frac{s_\alpha}{c_\beta}, \frac{c_\alpha}{c_\beta}, t_\beta \right\},$$

$$\lambda_{\chi_k^+}^{\{H^0, H^0, A^0\}} = \frac{\sqrt{2} M_W}{m_{\chi_k^+}} (U_{k1} V_{k2} \{c_\alpha, s_\alpha, -c_\beta\} + U_{k2} V_{k1} \{-s_\alpha, c_\alpha, -s_\beta\}).$$

$$\lambda_{\bar{l}}^{\{H^0, H^0\}} = \frac{2m_{\bar{l}}}{m_{\bar{l}}^2 c_\beta} (+\mu^* \{s_\alpha, -c_\alpha\} + A_t \{c_\alpha, s_\alpha\}) (U_{\bar{l}1}^t)^* U_{\bar{l}2}^t,$$

$$\lambda_{\bar{b}}^{\{H^0, H^0\}} = \frac{2m_{\bar{b}}}{m_{\bar{b}}^2 c_\beta} (-\mu^* \{c_\alpha, s_\alpha\} + A_b \{-s_\alpha, c_\alpha\}) (U_{\bar{b}1}^b)^* U_{\bar{b}2}^b,$$

$$\lambda_{\bar{\tau}}^{\{H^0, H^0\}} = \frac{2m_{\bar{\tau}}}{m_{\bar{\tau}}^2 c_\beta} (-\mu^* \{c_\alpha, s_\alpha\} + A_\tau \{-s_\alpha, c_\alpha\}) (U_{\bar{\tau}1}^\tau)^* U_{\bar{\tau}2}^\tau.$$

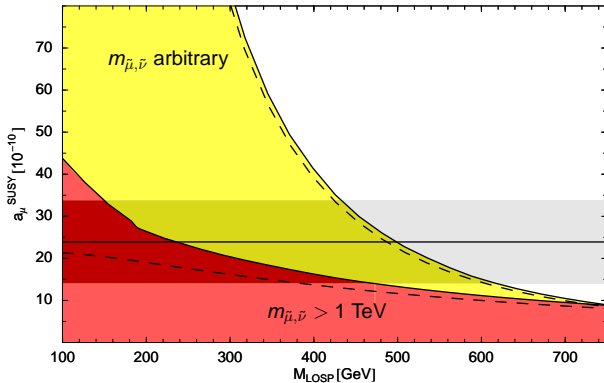
$$f_{\text{PS}}(z) = \frac{2z}{y} \left[\text{Li}_2\left(1 - \frac{1-y}{2z}\right) - \text{Li}_2\left(1 - \frac{1+y}{2z}\right) \right]$$

$$f_S(z) = (2z-1)f_{\text{PS}}(z) - 2z(2+\log z),$$

$$f_{\bar{l}}(z) = \frac{z}{2} [2 + \log z - f_{\text{PS}}(z)].$$

Numerical result

$\tan\beta = 50$, all parameters < 3 TeV

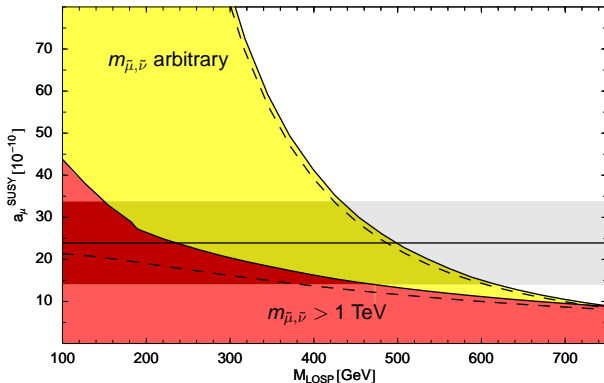


- typically:
 $13 \times 10^{-10} \tan\beta$
 $\times \left(\frac{100 \text{ GeV}}{M_{\text{SUSY}}} \right)^2 \text{sign}(\mu)$

SUSY contributions in the observed range for **low** M_{SUSY} !

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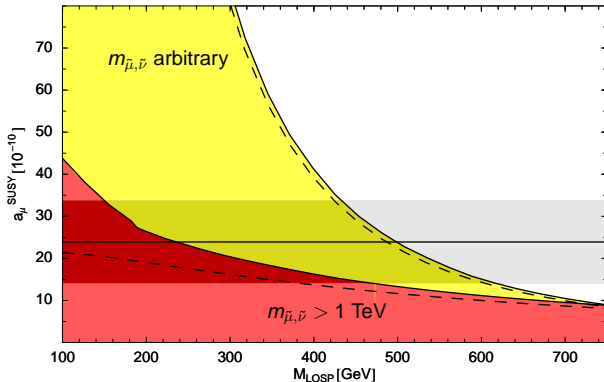
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- typically:
 $13 \times 10^{-10} \tan\beta$
 $\times \left(\frac{100 \text{ GeV}}{M_{\text{SUSY}}}\right)^2 \text{sign}(\mu)$
- possible enhancement for large μ
- suppression for large $m_{\tilde{\mu}, \tilde{\nu}} \leftrightarrow$
 2-Loop important

SUSY contributions in the observed range for low M_{SUSY} !

Conclusions

Two precision observables: M_W and $(g-2)_\mu$

- current measurements sensitive to 2-Loop SM and SUSY effects
- 2-Loop SUSY contributions:
 - M_W : Computer code for M_W^{MSSM} [Haestier, Heinemeyer, DS, Weiglein '05]
[Heinemeyer, DS, Weiglein '03,'04]
[Heinemeyer, Hollik, DS, Weber, Weiglein '06]
 - $(g-2)_\mu$: 1-,2-Loop contributions easy to implement
- SUSY with low mass scale $\sim 200 \dots 600$ GeV fits very well
- SM cannot be excluded at the moment

Future, more precise measurements very important and promising!