

Patterns of supersymmetry breaking in moduli-mixing racetrack model

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Based on articles

- [Phys.Rev.D73 \(2006\) 046005](#) ([hep-th/0511160](#))
- [Nucl.Phys.B742 \(2006\) 187](#) ([hep-th/0512232](#))
- [hep-th/0606095](#)

with Tetsutaro Higaki and Tatsuo Kobayashi

I. Introduction

Candidate of unified theory of gravity, gauge and matter fields:

(higher-dimensional) SUGRA, **superstring/M-theory**

The 4D effective theory has many moduli superfields $\Phi^I = (\phi, \chi, F^\phi)^I$.

$$\langle \phi \rangle \longrightarrow \begin{cases} M_{Pl} & : \text{ Planck scale,} \\ g_a & : \text{ gauge couplings,} \\ Y_{ijk} & : \text{ Yukawa couplings,} \\ \vdots & \end{cases}$$

$$\langle F^\phi \rangle \longrightarrow \begin{cases} M_a & : \text{ gaugino masses,} \\ m_i & : \text{ squark, slepton masses,} \\ A_{ijk} & : \text{ A-terms,} \\ \vdots & \end{cases}$$

Thus, **moduli stabilization and its effect on SUSY breaking** is quite relevant to **particle phenomenology/cosmology**.

Sources of moduli potential (in Type IIB $O3/O7$ on CY_3 with $h_{1,1} = 1$)

Gaugino condensation: $SU(N_a)$ SYM, $1/g_a^2 = \text{Re } f_a$

$$W_{GC} = N_a e^{-8\pi^2 f_a / N_a} \longleftarrow \begin{cases} f_{D7} = T & : \text{Kähler modulus} \\ f_{D3} = S & : \text{dilaton} \end{cases}$$

Three-form flux: $G_3 = F_3 - 2\pi i S H_3$

$$\begin{aligned} W_{flux} &= \int_{CY_3} G_3 \wedge \Omega && (\Omega : \text{holomorphic three-form}) \\ &= f^{RR}(U^\alpha) + S f^{NS}(U^\alpha) && \longleftarrow \begin{cases} S & : \text{dilaton} \\ U^\alpha & : \text{complex structure moduli} \end{cases} \\ &&& f^{RR,NS}(U^\alpha) : \text{functions determined by flux} \end{aligned}$$

Gukov, Vafa, Witten, Nucl.Phys.B584 (2000) 69

Similarly, moduli potential can arise also in other models (e.g. IIA, hetero)

Two stage stabilization (KKLT-type)

Kachru, Kallosh, Linde, Trivedi, Phys.Rev.D68 (2003) 046005

Choi, Falkowski, Nilles, Olechowski, Pokorski, JHEP 0411 (2004) 076, ...

H.A., Higaki, Kobayashi, hep-th/0606095

By fine-tuning flux such that

$$w_0 = \langle W_{flux} \rangle \ll 1 \quad (M_{Pl} = 1 \text{ unit})$$

while keeping $\langle \partial_S \partial_{U^\alpha} W_{flux} \rangle, \langle \partial_{U^\alpha} \partial_{U^\beta} W_{flux} \rangle \sim 1$

we find ...

- Moduli S and U^α can be stabilized by flux and receive heavy mass ~ 1
- Remaining T is stabilized by GC with mass $\sim m_{3/2} \sim w_0 \sim \langle W_{GC} \rangle \ll 1$

We first integrate out S, U^α , and derive the effective action of T .

II. 4D effective SUGRA for single modulus

Stabilization at SUSY AdS minimum: $F^T = -e^{K/2} K^{\bar{T}T} (W_T + K_T W) = 0$

$$V = K_{I\bar{J}} F^I \bar{F}^{\bar{J}} - 3e^K |W|^2 \quad \rightarrow \quad -3m_{3/2}^2 < 0$$

Flux + GC: $K = -3 \ln(T + \bar{T}), \quad W = w_0 - Ae^{-aT}$

$$at \simeq \ln(A/w_0) \sim 4\pi^2, \quad \sigma_t = 0, \quad (T = t + i\sigma_t)$$

$$A \approx 1, \quad m_{3/2} \simeq w_0 \sim 10^{-14} \quad (\text{fine-tuning})$$

Racetrack : $K = -3 \ln(T + \bar{T}), \quad W = Ae^{-aT} - Be^{-bT}$

$$(a - b)t \simeq \ln(aA/bB) \quad \longrightarrow \quad at \approx bt \approx 4\pi^2$$

$$m_{3/2} \simeq Ae^{-aT} \sim 10^{-14}, \quad |a - b| \equiv \frac{a + b}{\ln \frac{M_{Pl}}{m_{3/2}}} \quad (\text{fine-tuning})$$

Uplifting AdS to dS (Minkowski) (KKLT model)

Kachru, Kallosh, Linde and Trivedi, Phys.Rev.D68 (2003) 046005

$$V_{\text{lift}} = D(T + \bar{T})^{n_P-2} \quad \longrightarrow \quad V + V_{\text{lift}} = V_{\text{dS}} \geq 0$$

At the linear order of δ^t , $t = t_{SUSY}(1 + \delta^t)$, $F^T|_{t=t_{SUSY}} = 0$

Choi, Falkowski, Nilles, Olechowski, Nucl.Phys.B718 (2005) 113

$$\delta^t \simeq \frac{1 - n_P}{2(at)^2} \quad : (n_P = 0 \text{ for } \bar{D}3)$$

SUSY breaking order parameters

flux+GC : $\frac{F^C}{C_0} \simeq m_{3/2} \approx w_0, \quad \frac{F^T}{T + \bar{T}} \simeq \frac{2 - n_P}{2} \frac{1}{at} m_{3/2}$

Racetrack : $\frac{F^C}{C_0} \simeq m_{3/2} \approx Ae^{-at}, \quad \frac{F^T}{T + \bar{T}} \simeq \frac{3(2 - n_P)}{4} \frac{1}{at bt} m_{3/2}$

$$\alpha \equiv \frac{1}{4\pi^2} \frac{F^C}{C_0} \frac{T + \bar{T}}{F^T} \simeq \begin{cases} \frac{at}{\ln(M_{Pl}/m_{3/2})} \frac{2}{2 - n_P} \sim 1 & : \text{flux+GC} \\ \frac{at bt}{\ln(M_{Pl}/m_{3/2})} \frac{4}{3(2 - n_P)} \sim 4\pi^2 & : \text{racetrack} \end{cases}$$

III. Moduli-mixing racetrack model

Moduli-mixing gauge coupling $\text{Re } f_a = 1/g_a^2, \quad S = s + i\sigma_s, \quad T = t + i\sigma_t$

Type IIB $O3/O7$, magnetized D -branes:

Lüst, Mayr, Richter, Stieberger, Nucl.Phys. B696 (2004) 205

$$\begin{aligned} f_{mD7} &= |m_{D7}|S + |w_{D7}|T \\ f_{mD9} &= m_{D9}S - w_{D9}T \end{aligned} \quad \left\{ \begin{array}{l} m_{D7,D9} : \text{ magnetic flux number} \\ w_{D7,D9} : \text{ winding number} \end{array} \right.$$

Heterotic (M-)theory:

Choi, Kim, Phys.Lett.B165 (1985) 71, ...

Lukas, Ovrut, Waldram, Nucl.Phys. B532 (1998) 43

$$f_H = S - \beta T + f_{M5}, \quad \beta \sim \frac{1}{16\pi^4} \int_{CY} J \wedge \left(\text{tr } F_2^2 - \frac{1}{2} \text{tr } R^2 \right)$$

$SU(N_a)$ gaugino condensation:

$$W_{GC} = N_a e^{-8\pi^2 f_a / N_a} \longrightarrow \text{moduli-mixing superpotential}$$

Single light modulus

H.A., Higaki, Kobayashi, Phys.Rev.D73 (2006) 046005

Assuming $\langle \partial_S \partial_{U^\alpha} W_{flux} \rangle \neq 0$, $\langle W_{flux} \rangle = 0$ (S is stabilized at high scale)

$$K = -\ln \langle S + \bar{S} \rangle - 3 \ln(T + \bar{T})$$

$$W = A e^{-a(m_a \langle S \rangle \pm w_a T)} \pm B e^{-b(m_b \langle S \rangle \pm w_b T)}$$

Typical models: $A' = A e^{-am_a \langle S \rangle}$, $B' = B e^{-bm_b \langle S \rangle}$

$$W_{D3+mD7} = A' - B' e^{-bw_b T} \quad : \quad 0 < \alpha \lesssim 1$$

$$W_{D3+mD9} = A' + B' e^{+bw_b T} \quad : \quad -1 \lesssim \alpha < 0$$

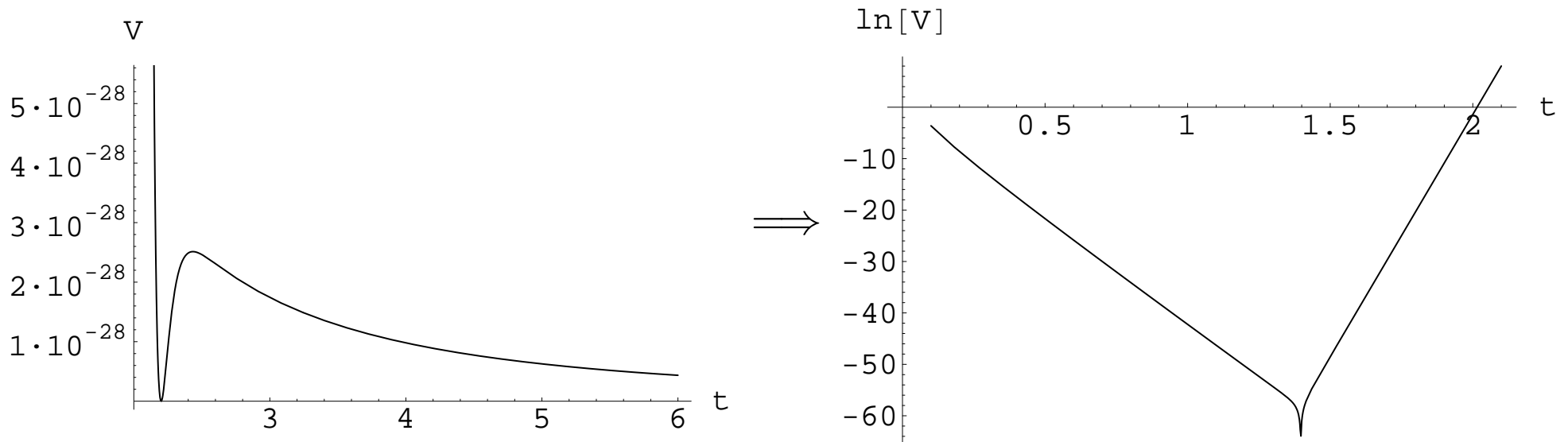
$$W_{mD7+mD7'} = A' e^{-aw_a T} - B' e^{-bw_b T} \quad : \quad 1 < \alpha \lesssim 4\pi^2$$

$$W_{mD7+mD9} = A' e^{-aw_a T} + B' e^{+bw_b T} \quad : \quad -4\pi^2 \lesssim \alpha < -1$$

$$\alpha = \frac{1}{4\pi^2} \frac{F^C T + \bar{T}}{C_0 F^T}$$

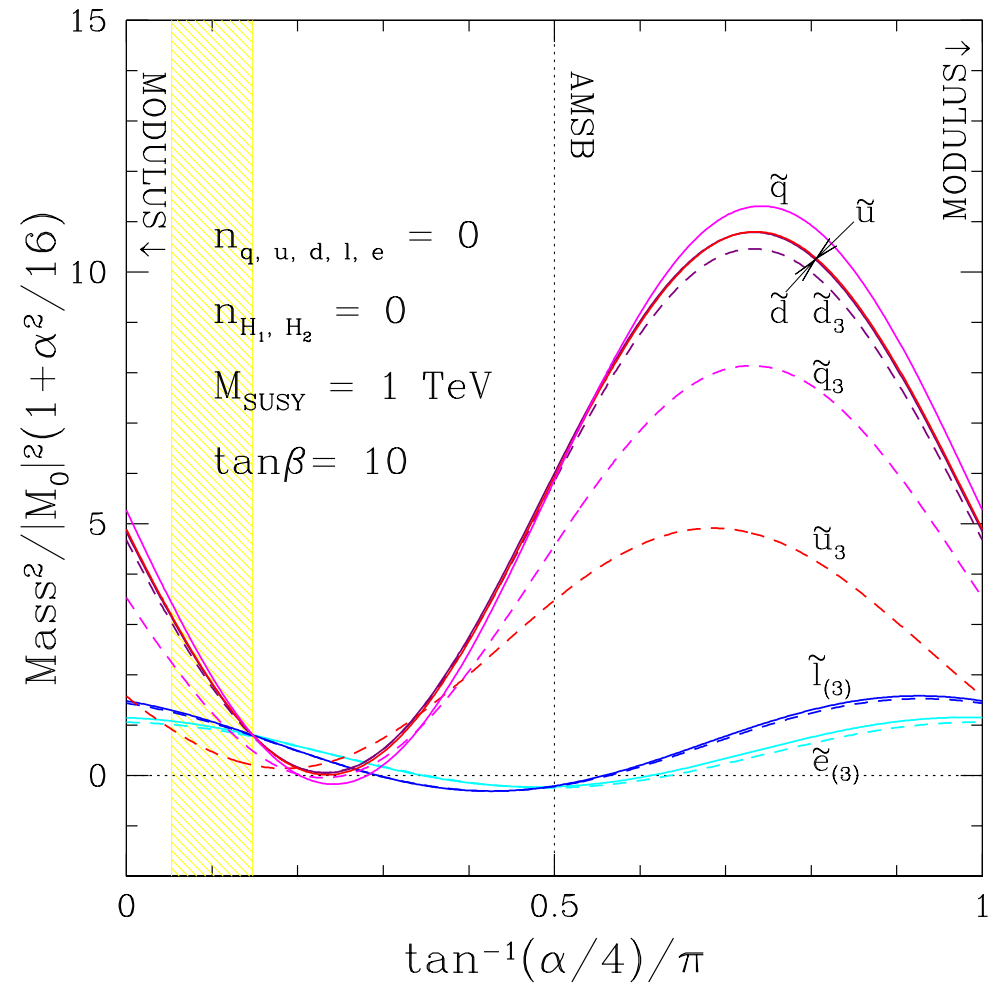
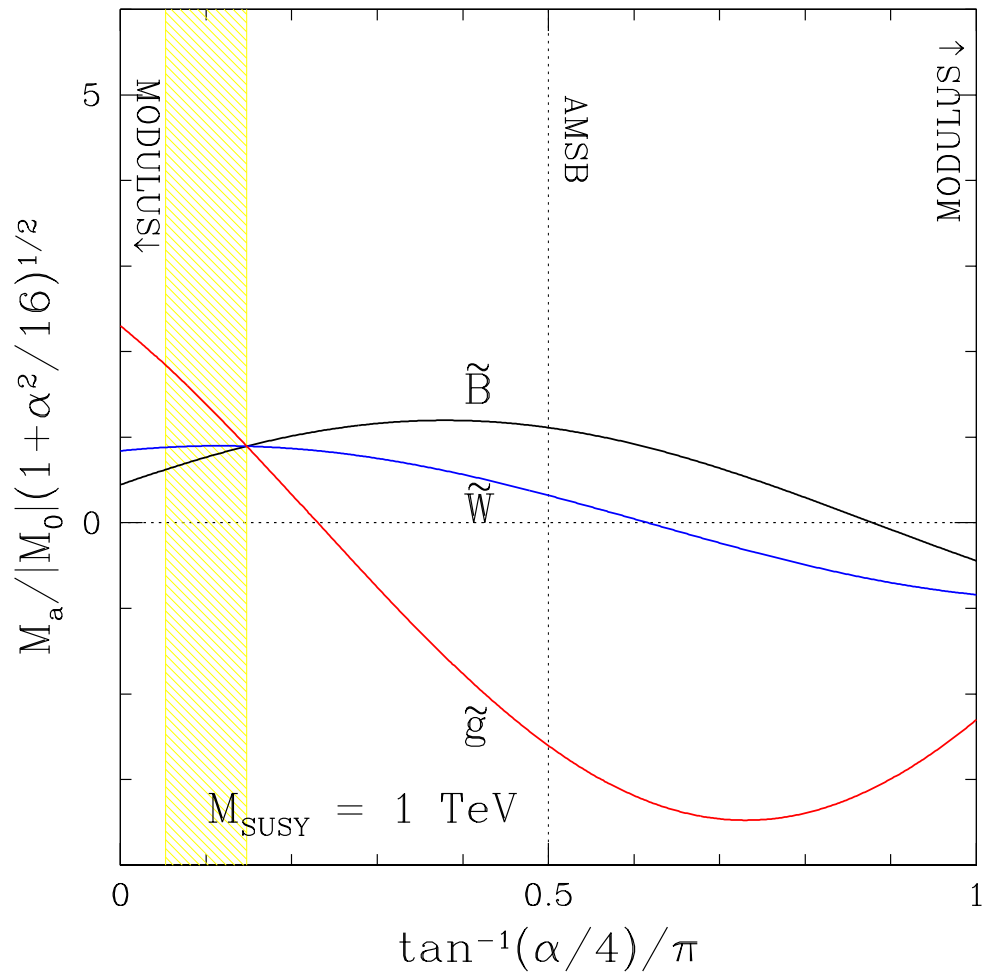
Implications on SUSY phenomenology/cosmology

- Fine-tuning $w_0 \ll 1$ is not necessary if $\langle W_{flux} \rangle = 0$ and Ae^{-aS} exists in W_{GC} .
- Destabilization/overshooting problem may be avoided in some case.



- Modulus/anomaly ratio of SUSY breaking mediation, α can take various value depending on m , w and $\langle S \rangle = -\langle \partial_U f^{RR}(U) / \partial_U f^{NS}(U) \rangle$.

Gaugino and sfermion masses at TeV scale



Choi, Jeong, Okumura, JHEP09 (2005) 039

Two light moduli

H.A., Higaki, Kobayashi, Nucl.Phys.B742 (2006) 187

Assuming $\langle \partial_S \partial_{U^\alpha} W_{flux} \rangle = 0$, $\langle W_{flux} \rangle = 0$ (S also remains light)

$$K = -n_S \ln(S + \bar{S}) - n_T \ln(T + \bar{T})$$

$$W = Ae^{-a(S-\alpha T)} - Be^{-b(S+\beta T)}$$

$a, b, \alpha, \beta, A, B, n_S, n_T$: parameters of effective SUGRA

SUSY stationary point: $F^S, F^T = 0$ for $|s_{SUSY}| \gg n_S/2a, n_S/2b$

$$s_{SUSY} \sim \frac{1}{b-a} \ln \frac{bB}{aA}, \quad t_{SUSY} \simeq \frac{n_T(b-a)}{2ab(\alpha+\beta)}, \quad (W_{SUSY} \neq 0)$$

Negative vacuum energy: $V_{SUSY} = -3(m_{3/2}^{SUSY})^2 < 0$

$$m_{3/2}^{SUSY} \sim \frac{b-a}{b} \left(\frac{aA}{bB} \right)^{\frac{a}{b-a}} \frac{A}{(2s_{SUSY})^{n_S/2} (2t_{SUSY})^{n_T/2}}$$

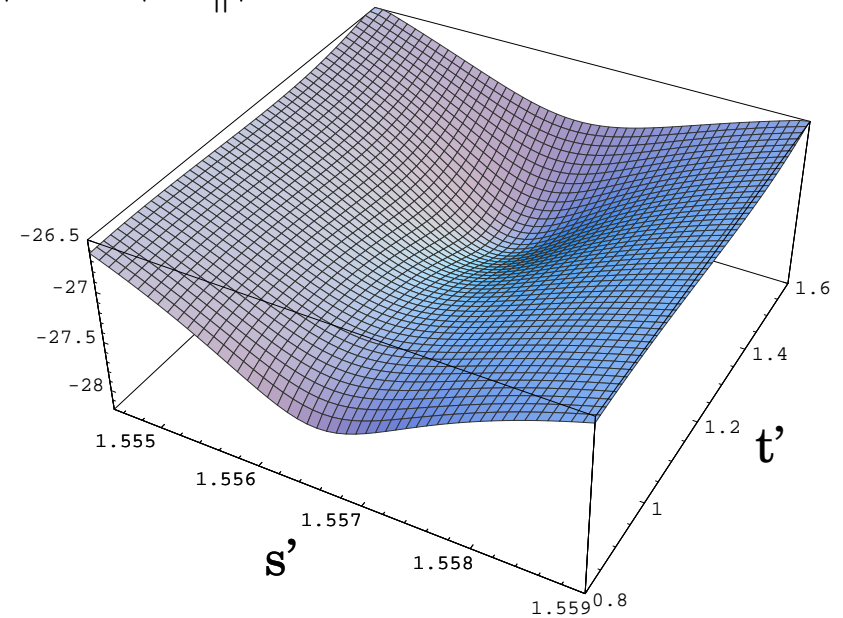
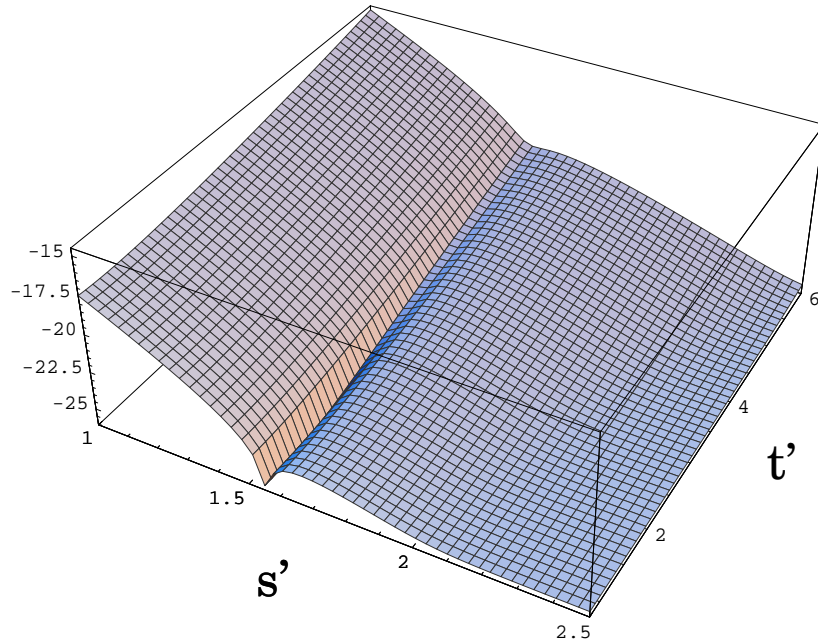
For $s_{SUSY}, t_{SUSY} > 1$ with $a, b \gg 1$ and $A, B \sim \mathcal{O}(1)$

$$b - a \sim \mathcal{O}(1), \quad B/A > a/b, \quad \alpha, \beta \sim \mathcal{O}(1/ab),$$

Stability of SUSY point: mass eigenvalues $(m_{\perp}^2, m_{\parallel}^2)$

$$\begin{cases} m_{\perp}^2 \sim \frac{128}{n_S^2} \left(\ln \frac{bB}{aA} \right)^4 \frac{(ab)^2}{(b-a)^4} (m_{3/2}^{SUSY})^2 > 0 \\ m_{\parallel}^2 \sim -(m_{3/2}^{SUSY})^2 < 0 \end{cases} \longrightarrow \text{saddle point}$$

$$|m_{\perp}^2| \gg |m_{\parallel}^2| \longrightarrow \text{sharp racetrack}$$



SUSY breaking local minimum $t_{SB} = t_{SUSY}(1 + \delta_{SB}^t)$, $s_{SB} = s_{SUSY}(1 + \delta_{SB}^s)$

$$\delta_{SB}^t = \frac{\sqrt{n_T + 1} - 1}{n_T} \lesssim \mathcal{O}(1), \quad \delta_{SB}^s = -\frac{a\alpha + b\beta}{\ln(bB/aA)} \delta_{SB}^t$$

Order parameters

$$m_{3/2} \simeq \frac{e^{-a(\delta_{SB}^s - \alpha\delta_{SB}^t)}}{(1 + \delta_{SB}^s)^{n_S/2} (1 + \delta_{SB}^t)^{n_T/2}} m_{3/2}^{SUSY},$$

$$\frac{F^T}{T + \bar{T}} \simeq -\delta_{SB}^t m_{3/2}, \quad \frac{F^S}{S + \bar{S}} \sim \frac{n_S}{as} \frac{F^T}{T + \bar{T}} \ll \frac{F^T}{T + \bar{T}}$$

Vacuum energy

$$V_{SB} = n_S \left| \frac{F^S}{S + \bar{S}} \right|^2 + n_T \left| \frac{F^T}{T + \bar{T}} \right|^2 - 3m_{3/2}^2 < -2m_{3/2}^2$$

AdS minimum with modulus-dominated SUSY breaking

$$\Downarrow \quad V_{\text{lift}} = D e^{2K/3} (T + \bar{T})^{n_P} (S + \bar{S})^{m_P}$$

Minkowski minimum with modulus-dominated SUSY breaking $\alpha \sim \frac{1}{4\pi^2} \ll 1$

Numerical results

s_{SUSY}	t_{SUSY}	V_{SUSY}	$(m_{\perp}^{SUSY}/m_{3/2}^{SUSY})^2$	$ m_{\parallel}^{SUSY}/m_{3/2}^{SUSY} ^2$
1.41	1.18	-1.79×10^{-26}	1.11×10^6	4.00
s_{SB}	t_{SB}	V_{SB}	$(m_{\perp}^{SB}/m_{3/2})^2$	$(m_{\parallel}^{SB}/m_{3/2})^2$
1.36	1.57	-1.82×10^{-26}	9.55×10^5	7.09
s_{dS}	t_{dS}	V_{dS}	$(m_{\perp}^{dS}/m_{3/2}^{dS})^2$	$(m_{\parallel}^{dS}/m_{3/2}^{dS})^2$
1.30	1.96	$+1.20 \times 10^{-34}$	8.16×10^5	4.15×10
AdS	$F^S/(S + \bar{S})$	$F^T/(T + \bar{T})$	$m_{3/2}$	$m_{3/2}^{SUSY}$
—	4.38×10^{-15}	-2.77×10^{-14}	8.28×10^{-14}	7.72×10^{-14}
dS	$F^S/(S + \bar{S})$	$F^T/(T + \bar{T})$	$m_{3/2}^{dS}$	D
—	1.34×10^{-14}	-6.49×10^{-14}	9.78×10^{-14}	2.45×10^{-25}

IV. Summary

Moduli-mixing racetrack model

$$W = Ae^{-af_a} - Be^{-bf_b}, \quad f_{a,b} = m_{a,b}S + w_{a,b}T$$

Single light modulus $S \rightarrow \langle S \rangle, T$

Modulus/anomaly ratio of SUSY breaking mediation α can take various values depending on m, w and $\langle S \rangle$

(c.f. $\alpha = 1$ in KKLT model without moduli mixing)

Two light moduli S, T

SUSY point is saddle point \rightarrow SUSY breaking AdS local minimum

Modulus-dominated SUSY breaking $\alpha \ll 1$

(c.f. $\alpha \gg 1$ in racetrack model without moduli mixing)